

**Derivation of A(N) in Eq. 16 (from the main text)**

The (normalized) probability mass function (PMF) for a power law distribution of cluster sizes  $s = 1, 2, \dots, N$  is

$$P(s) = \frac{s^\alpha}{1 + 2^\alpha + \dots + N^\alpha}. \quad (1)$$

For the PMF of rescaled sizes,  $z = s/N$ , we write

$$P_z(z) = A(N)z^\alpha, \quad (2)$$

where  $A(N)$  denotes the normalization factor that depends on the system size  $N$ . From Eq. 2 and with the property

$$\sum_{s=1}^N P_z(s/N) = 1,$$

one obtains

$$A(N) \left[ \left( \frac{1}{N} \right)^\alpha + \left( \frac{2}{N} \right)^\alpha + \dots + \left( \frac{N}{N} \right)^\alpha \right] = 1,$$

and thus

$$A(N) = \frac{N^\alpha}{1 + 2^\alpha + \dots + N^\alpha}, \quad (3)$$

which is equivalent to Eq. 16 from the main text.

Analogously, for the exponential distribution  $P_z(z) = A(N)e^{-\lambda z}$  with  $z = s/N$  ( $s = 1, \dots, N$ ), one obtains

$$A(N) = \frac{1}{e^{-\lambda/N} + e^{-2\lambda/N} + \dots + e^{-N\lambda/N}}. \quad (4)$$

For the interested reader, we also provide the results for the continuous power law distribution. The probability density function for cluster sizes  $b \leq s \leq N$  and exponent  $\alpha < -1$  is given by

$$P(s) = \frac{\alpha + 1}{N^{\alpha+1} - b^{\alpha+1}} s^\alpha, \quad (5)$$

where  $b$  denotes the lower bound for  $s$ . The normalized probability density for a rescaled system of size  $N$  is defined as

$$P_z(z) = A(N)z^\alpha. \quad (6)$$

Since  $s$  ranges from  $b$  to  $N$ , the rescaled size  $z$  ranges from  $b/N$  to 1. This gives the following constraint for  $P_z(z)$ :

$$\int_{b/N}^1 P_z(z) dz = 1.$$

From this and with  $\alpha < -1$ , it follows that

$$\int_{b/N}^1 A(N)z^\alpha dz = \left[ \frac{A(N)}{\alpha + 1} z^{\alpha+1} \right]_{b/N}^1 = \frac{A(N)}{\alpha + 1} [1 - (b/N)^{\alpha+1}] = 1,$$

which yields

$$A(N) = \frac{\alpha + 1}{1 - (b/N)^{\alpha+1}}. \quad (7)$$

### Collapse of cluster size distributions

Here, we show that the transformation  $z = s/N$  and a proper rescaling of  $P(s)$ ,  $s = 1, 2, \dots, N$ , results in a collapse of power law distributions for different system sizes  $N$ . From  $z = s/N$  and Eqs. 1 to 3 it follows that  $P_z(z) = P(s)$ . Dividing  $P_z(z)$  (Eq. 2) by  $A(N)$  (Eq. 3) gives  $z^\alpha$ , which is independent of  $N$ . The rescaling for the discrete power law distribution is therefore given by  $z = s/N$  and  $P_z(z)/A(N) = P(s)/A(N)$ . Consequently, the rescaled probability

$$\frac{P(s)}{A(N)} = \frac{s^\alpha}{1 + 2^\alpha + \dots + N^\alpha} \frac{1 + 2^\alpha + \dots + N^\alpha}{N^\alpha} = \frac{s^\alpha}{N^\alpha}$$

is the same for a rescaled system with  $z = s/N = (ks)/(kN)$ , that is, for system size  $kN$  and cluster size  $ks$ :

$$\frac{P(ks)}{A(kN)} = \frac{(ks)^\alpha}{1 + 2^\alpha + \dots + (kN)^\alpha} \frac{1 + 2^\alpha + \dots + (kN)^\alpha}{(kN)^\alpha} = \frac{(ks)^\alpha}{(kN)^\alpha} = \frac{s^\alpha}{N^\alpha}.$$

Therefore, a power law with exponent  $\alpha$  results in a collapse of  $P(s)/A(N)$  for different  $N$ .

For the continuous power law distribution, the change in variables,  $z = s/N$ , yields  $P_z(z) = P(s)N$ . The proper rescaling of the probability density is therefore given by  $z = s/N$  and  $P_z(z)/A(N) = P(s)N/A(N)$ . Rescaling of the probability density  $P(s) = (\alpha + 1)/(N^{\alpha+1} - b^{\alpha+1})s^\alpha$ ,  $b \leq s \leq N$  (Eq. 5) for system size  $N$  and cluster size  $s$  yields

$$\frac{P(s)N}{A(N)} = \frac{\alpha + 1}{N^{\alpha+1} - b^{\alpha+1}} s^\alpha N \frac{(N^{\alpha+1} - b^{\alpha+1})/N^{\alpha+1}}{\alpha + 1} = \frac{s^\alpha}{N^\alpha},$$

which is equal to the value one obtains for a rescaled system with  $z = (ks)/(kN)$ , that is, for system size  $kN$  and cluster size  $ks$ :

$$\frac{P(ks)kN}{A(kN)} = \frac{\alpha + 1}{(kN)^{\alpha+1} - b^{\alpha+1}} (ks)^\alpha kN \frac{((kN)^{\alpha+1} - b^{\alpha+1})/(kN)^{\alpha+1}}{\alpha + 1} = \frac{s^\alpha}{N^\alpha}.$$

Therefore, a continuous power law with exponent  $\alpha$  results in a collapse of  $P(s)N/A(N)$  for different  $N$  (the same result can be obtained for a power law distribution with an upper bound  $\infty$ ).